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Classical mechanics

Central force; Definition and characteristics →

The properties of a particle of mass  $m$  moving in a particular type of force field, a central force field. Central forces are very important in physics and engineering. For example, the gravitational force of attraction between two point masses is a central force. The Coulomb force of attraction and repulsion between charged particles is a central force. Because of their importance they deserve special consideration. We begin a precise definition of central force, or central force field.

A force acting on a particle of mass  $m$  has the properties that:

\* The force is always directed from  $m$  toward or away, from a fixed point  $O$

\* The magnitude of the force only depends on the distance  $r$  from  $O$ .

Forces having these properties are called central forces. The particle is said to move in a central force field. The point  $O$  is referred to as the ~~central~~ center of force.

If  $f(r) < 0$  the force is said to be attractive towards  $O$ . If  $f(r) > 0$  the force is said to be repulsive from  $O$ .

Properties of a particle moving under the influence of a central force  $\rightarrow$  If a particle moves in a central force field then the following properties hold!

1. The path of the ~~centre~~ particle must be a plane curve, i.e. it must lie in a plane.
2. The angular momentum of the particle is conserved, i.e. it is constant in time.
3. The particle moves in such a way that the position vector (from the point  $O$ ) sweeps out equal

areas in equal times. In other words, the time rate of change in area is constant. This is referred to as the Law of Areas.

Reduction to the Equivalent one-body problem in terms of Lagrangian:-

Consider a monogenic system of two mass points,  $m_1$  and  $m_2$  (fig. 1), where the only forces are those due to an interaction potential  $U$ . Firstly we assume that  $U$  is any function of the vector between the two particles,  $\vec{r}_2 - \vec{r}_1$ , or of their relative velocity  $\dot{\vec{r}}_2 - \dot{\vec{r}}_1$ , or of any higher derivatives of  $\vec{r}_2 - \vec{r}_1$ . Such a system has six degrees of freedom and hence six independent generalized coordinates. If  $\vec{R}$  be the radius vector to the centre of mass, the Lagrangian will have the form;

$$L = T(\dot{\vec{R}}, \dot{\vec{r}}) - U(\vec{r}, \dot{\vec{r}}) \quad \text{--- (1)}$$

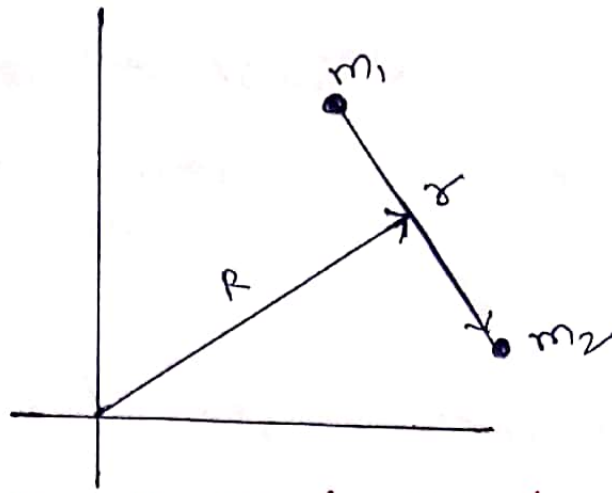


Fig.1 Coordinates for the two-body problem.

Now the kinetic energy  $T$  can be written as the sum of the kinetic energy of the motion of the centre of mass, and the kinetic energy of motion about the centre of mass  $T'$  :

$$T = \frac{1}{2} (m_1 + m_2) \dot{R}^2 + T'$$

with

$$T' = \frac{1}{2} m_1 \dot{r}_1'^2 + \frac{1}{2} m_2 \dot{r}_2'^2$$

Here  $\vec{r}_1'$  and  $\vec{r}_2'$  are the radii vectors of the two particles relative to the centre of mass and are related to  $\vec{r}$  by

$$\vec{r}_1' = - \frac{m_2}{m_1 + m_2} \vec{r}$$

$$\vec{r}_2' = \frac{m_1}{m_1 + m_2} \vec{r} \quad \text{————— (2)}$$

Substituting the value of  $\vec{r}_1$  &  $\vec{r}_2$ ,  $T'$  we get

$$T' = \frac{1}{2} \frac{m_1 m_2}{m_1 + m_2} \dot{\vec{r}}^2$$

and the total Lagrangian eq<sup>n</sup> (1) is

$$L = \frac{m_1 + m_2}{2} \dot{\vec{R}}^2 + \frac{1}{2} \frac{m_1 m_2}{m_1 + m_2} \dot{\vec{r}}^2 - U(\vec{r}, \vec{r}, \dots) \quad (3)$$

It is seen that the three coordinates  $\vec{R}$  are cyclic, so that the centre of mass is either at rest or moving uniformly.

The rest of the Lagrangian is exactly expected to be a fixed centre of force with a single particle at a distance  $r$  from it, having a mass

$$\mu = \frac{m_1 m_2}{m_1 + m_2} \quad (4)$$

where  $\mu$  is known as the reduced mass. Eq<sup>n</sup> (4) is written in the form

$$\frac{1}{\mu} = \frac{1}{m_1} + \frac{1}{m_2} \quad (5)$$

Thus the central force motion of two bodies about their centre of mass can always be reduced to an equivalent one-body problem.